

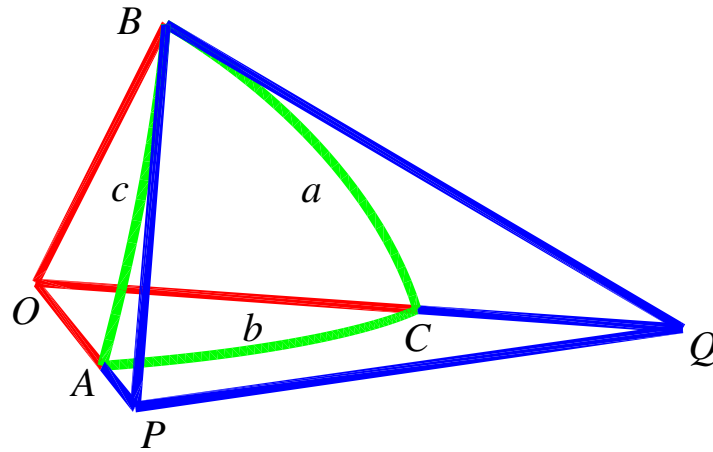
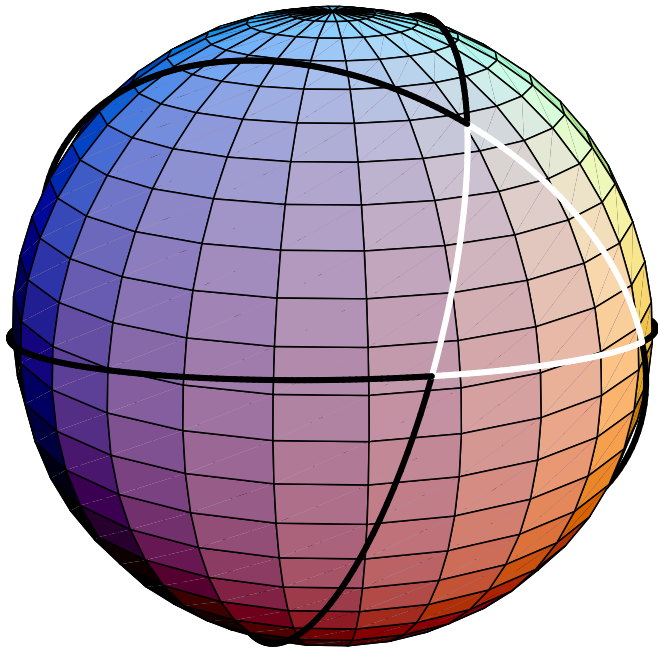
METR 5223: Atmospheric Radiation

# Spherical Trigonometry and Solar Radiation on a Sphere

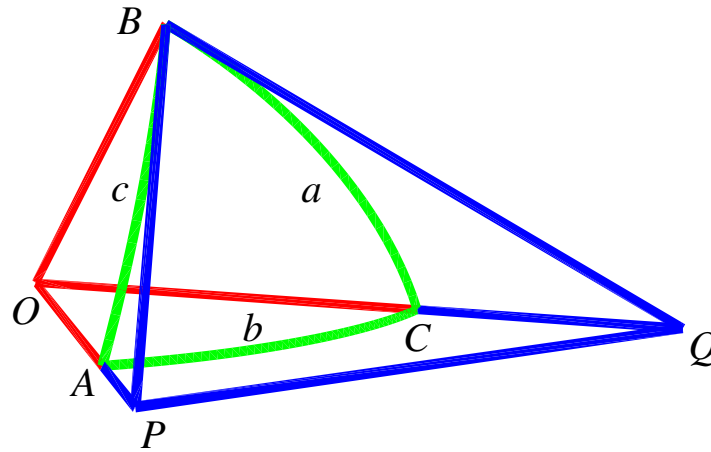
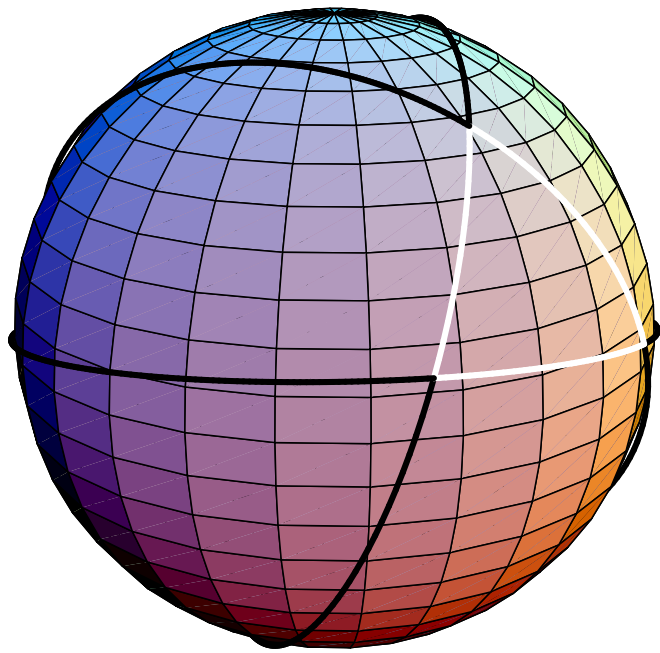
Lecture for Spring, 2009

Prof. Brian H. Fiedler

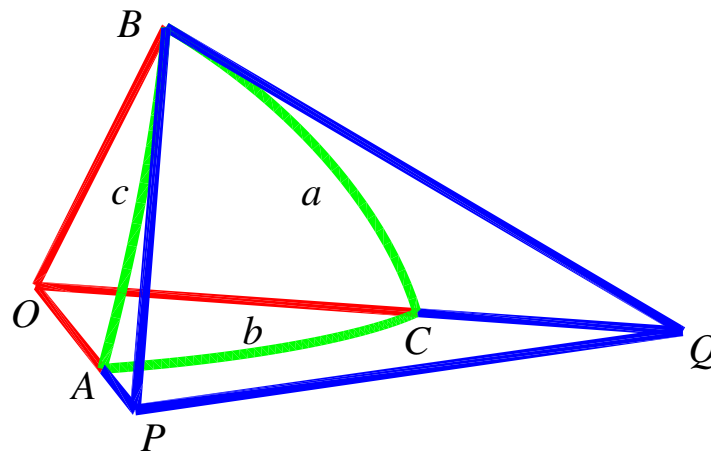
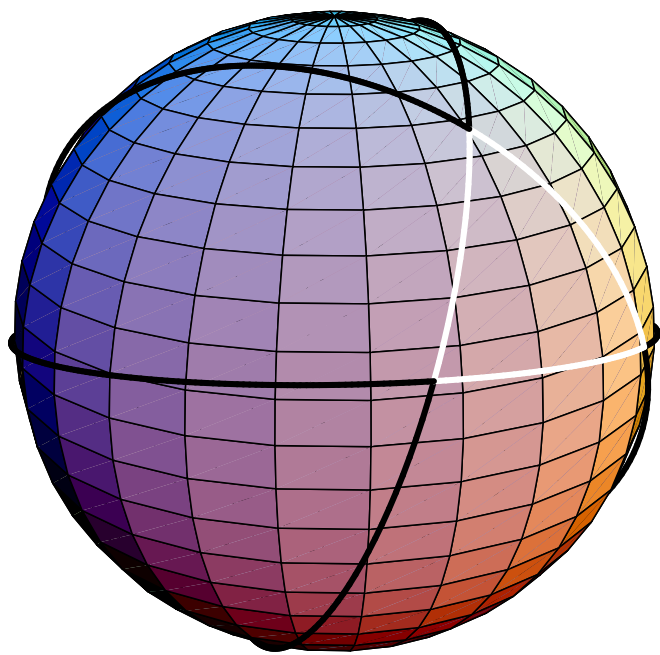
*School of Meteorology, University of Oklahoma*



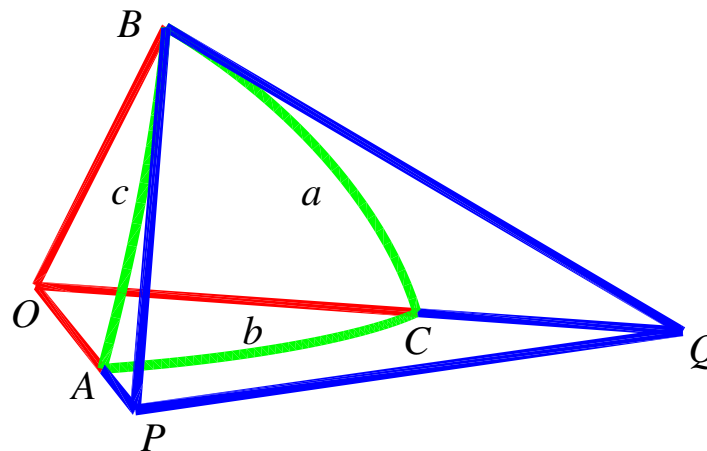
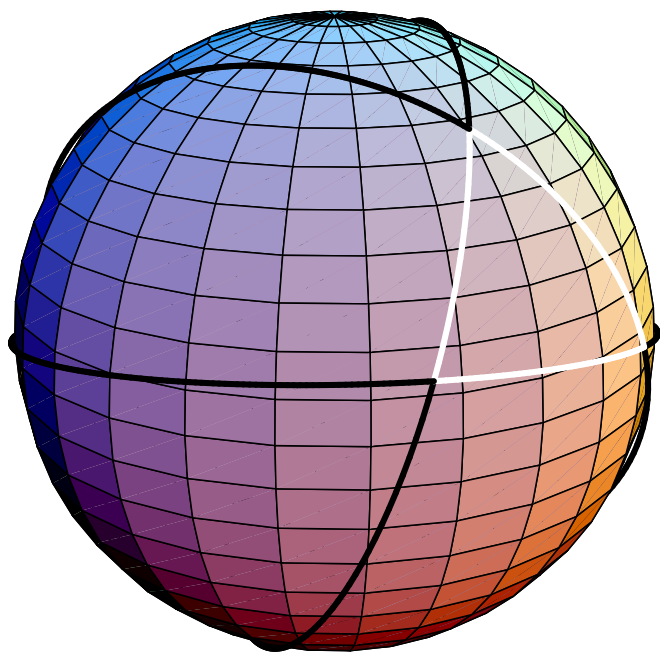
A great circle is formed by the intersection of a plane through the center of a sphere and the surface of the sphere. Three arcs of great circles can form a spherical triangle.



The points at the vertices of the spherical triangle are labeled  $A$ ,  $B$  and  $C$ . The angles at those points have measure  $A$ ,  $B$  and  $C$ . The sides opposite  $A$ ,  $B$  and  $C$  have arc angle  $a$ ,  $b$  and  $c$ .



The origin, or center of the sphere, is labeled  $O$ . Point  $P$  is at the intersection of a line through  $O$  and  $A$  and a plane tangent to the sphere at  $B$ . A similar definition applies to  $Q$ .

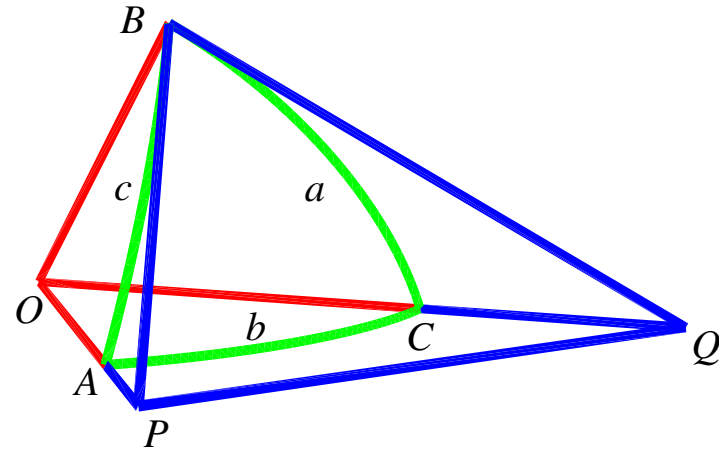


$A$  is measured in a plane tangent to the sphere.  $a$  is the arc angle measured in a plane through the center of the sphere.  $A$  is related to  $a$  by the *law of cosines* for a spherical triangle:

$$\cos a = \cos c \cos b + \sin c \sin b \cos A.$$

## First Proof

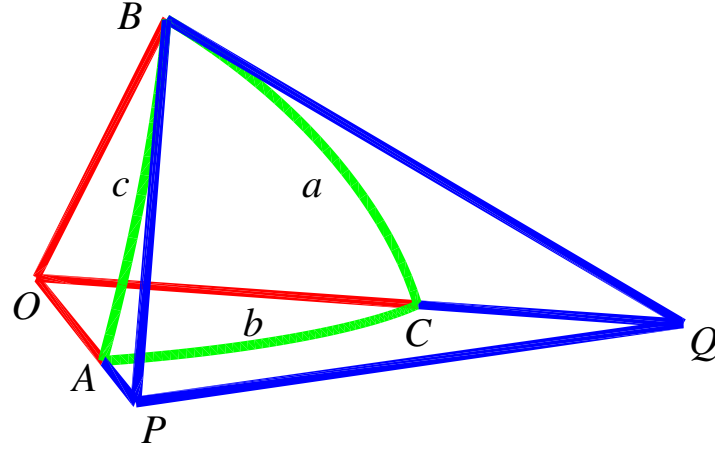
Angle  $A$  is the angle of intersection between planes  $OAB$  and  $OAC$ . This angle is the same as the angle between the normal vectors to those two planes.



$$\frac{\overrightarrow{OA}}{|\overrightarrow{OA}|} \equiv \hat{A} \quad \frac{\overrightarrow{OB}}{|\overrightarrow{OB}|} \equiv \hat{B} \quad \frac{\overrightarrow{OC}}{|\overrightarrow{OC}|} \equiv \hat{C}$$

The magnitude of  $\hat{A} \times \hat{B}$  is  $\sin c$  and the magnitude of  $\hat{A} \times \hat{C}$  is  $\sin b$ .

$$\frac{\hat{A} \times \hat{B}}{\sin c} \cdot \frac{\hat{A} \times \hat{C}}{\sin b} = \cos A$$



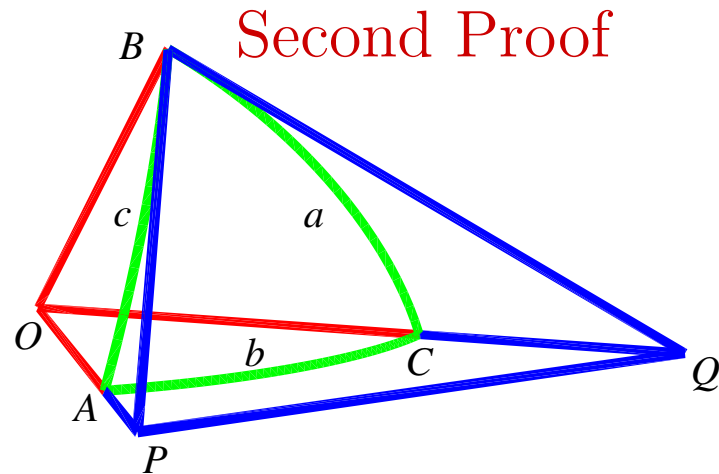
$$(\hat{A} \times \hat{B}) \cdot (\hat{A} \times \hat{C}) = \hat{B} \cdot \hat{C} \hat{A} \cdot \hat{A} - \hat{B} \cdot \hat{A} \hat{A} \cdot \hat{C}$$

$$(\hat{A} \times \hat{B}) \cdot (\hat{A} \times \hat{C}) = \cos a - \cos c \cos b$$

$$\cos a - \cos c \cos b = \cos A \sin c \sin b$$

or

$$\cos a = \cos c \cos b + \sin c \sin b \cos A.$$



Applying the law of cosines for plane triangles OPQ and BPQ we have:

$$(PQ)^2 = (OQ)^2 + (OP)^2 - 2(OQ)(OP) \cos b$$

and

$$(PQ)^2 = (BQ)^2 + (BP)^2 - 2(BQ)(BP) \cos B$$

Taking the difference of these equations gives

$$0 = (OQ)^2 - (BQ)^2 + (OP)^2 - (BP)^2 \\ - 2(OQ)(OP) \cos b + 2(BQ)(BP) \cos B$$

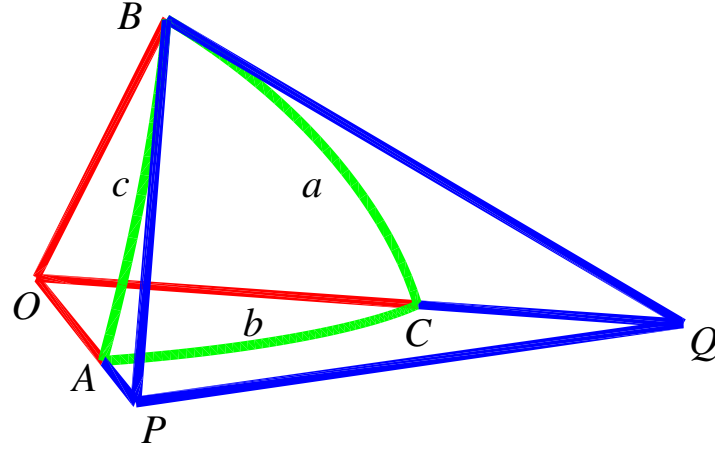
With the Pythagorean theorem giving

$$(OQ)^2 = (BQ)^2 + (OB)^2 \\ (OP)^2 = (BP)^2 + (OB)^2$$

$$0 = 2(OB)^2 - 2(OQ)(OP) \cos b + 2(BQ)(BP) \cos B$$

and thus

$$\cos b = \frac{OB}{OQ} \cdot \frac{OB}{OP} + \frac{BQ}{OQ} \cdot \frac{BP}{OP} \cos B$$

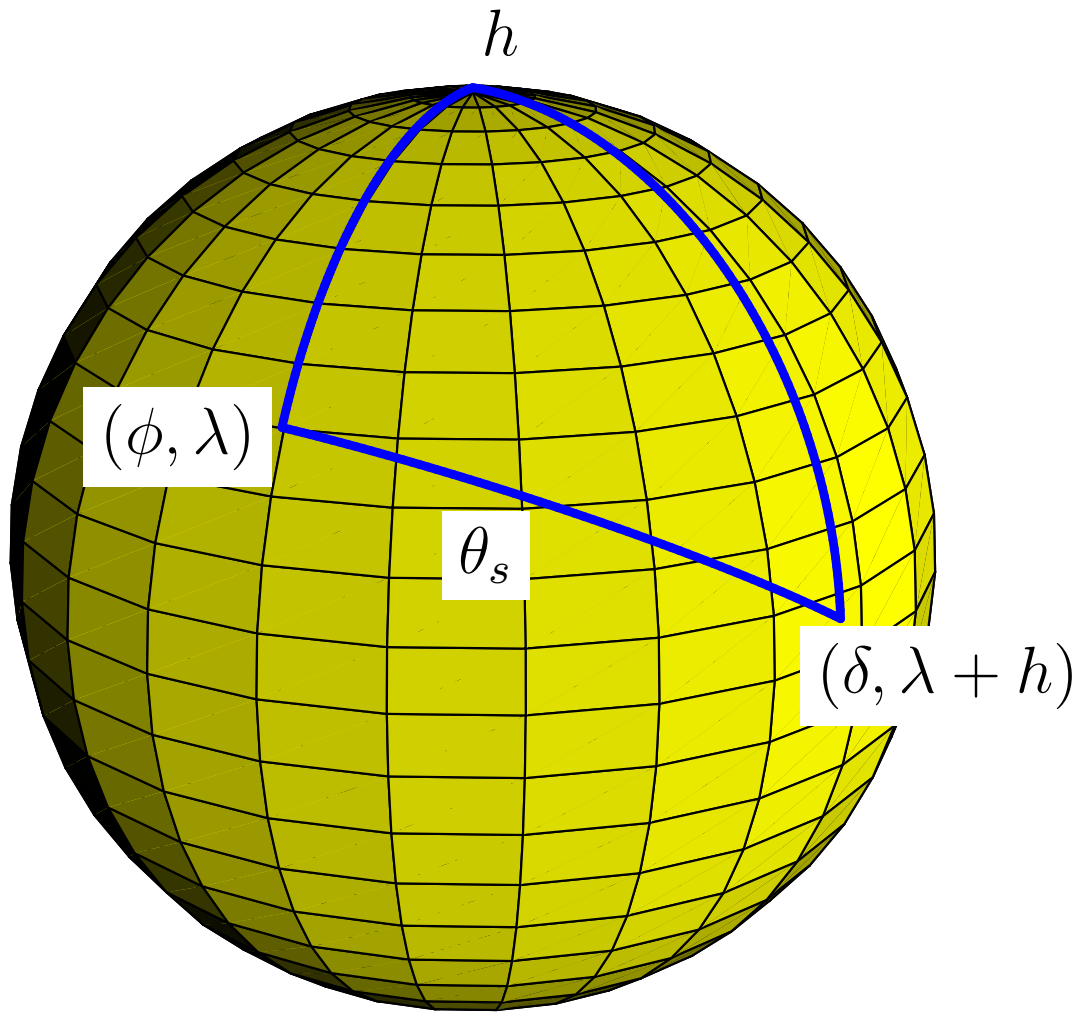


Noting the right triangles  $\triangle OBP$  and  $\triangle OBQ$ ,

$$\cos b = \frac{OB}{OQ} \cdot \frac{OB}{OP} + \frac{BQ}{OQ} \cdot \frac{BP}{OP} \cos B$$

is written as:

$$\cos b = \cos c \cos a + \sin c \sin a \cos B.$$



Calculate the zenith angle  $\theta_s$  at a lat-lon point  $(\phi, \lambda)$ . The subsolar point is at latitude  $\delta$  (the *solar declination angle*) and longitude  $\lambda + h$  where  $h$  is the *hour angle*.

$$\cos a = \cos c \cos b + \sin c \sin b \cos A.$$

$$\cos \theta_s = \cos \left( \frac{\pi}{2} - \phi \right) \cos \left( \frac{\pi}{2} - \delta \right) + \sin \left( \frac{\pi}{2} - \phi \right) \sin \left( \frac{\pi}{2} - \delta \right) \cos h$$

$$\cos \theta_s = \sin \phi \sin \delta + \cos \phi \cos \delta \cos h$$

The distance of Earth from the sun is  $R_E$ . The mean distance is  $R_o$ . The solar flux density into a plane tangent to the sphere, but at the top of the atmosphere:

$$Q = S_o \frac{R_o^2}{R_E^2} \cos \theta_s \text{ when } \cos \theta_s > 0$$

$$Q = 0 \text{ when } \cos \theta_s \leq 0$$

Sunrise occurs at hour angle  $h_o$ , (and sunset occurs at hour angle  $-h_o$ ), when  $\theta_s = \frac{\pi}{2}$ :

$$\sin \phi \sin \delta + \cos \phi \cos \delta \cos h_o = 0$$

The average value of  $Q$  over a day is:

$$\overline{Q}^{day} = \frac{\int_{h=\pi}^{h=-\pi} Q dh}{\int_{h=\pi}^{h=-\pi} dh}$$

Now  $\int_{h=\pi}^{h=-\pi} dh = -2\pi$  and, if  $\frac{R_o^2}{R_E^2}$  is nearly constant over the course of a day,

$$\begin{aligned} \int_{h=\pi}^{h=-\pi} Q dh &= \int_{h=h_o}^{h=-h_o} Q dh = S_o \frac{R_o^2}{R_E^2} \int_{h=h_o}^{h=-h_o} \cos \theta_s dh \\ &= S_o \frac{R_o^2}{R_E^2} [h \sin \phi \sin \delta + \cos \phi \cos \delta \sin h]_{h=h_o}^{h=-h_o} \\ &= -2S_o \frac{R_o^2}{R_E^2} [h_o \sin \phi \sin \delta + \cos \phi \cos \delta \sin h] \end{aligned}$$

Summary:

$$\overline{Q}^{day} = \frac{S_o}{\pi} \frac{R_o^2}{R_E^2} (h_o \sin \phi \sin \delta + \cos \phi \cos \delta \sin h_o)$$

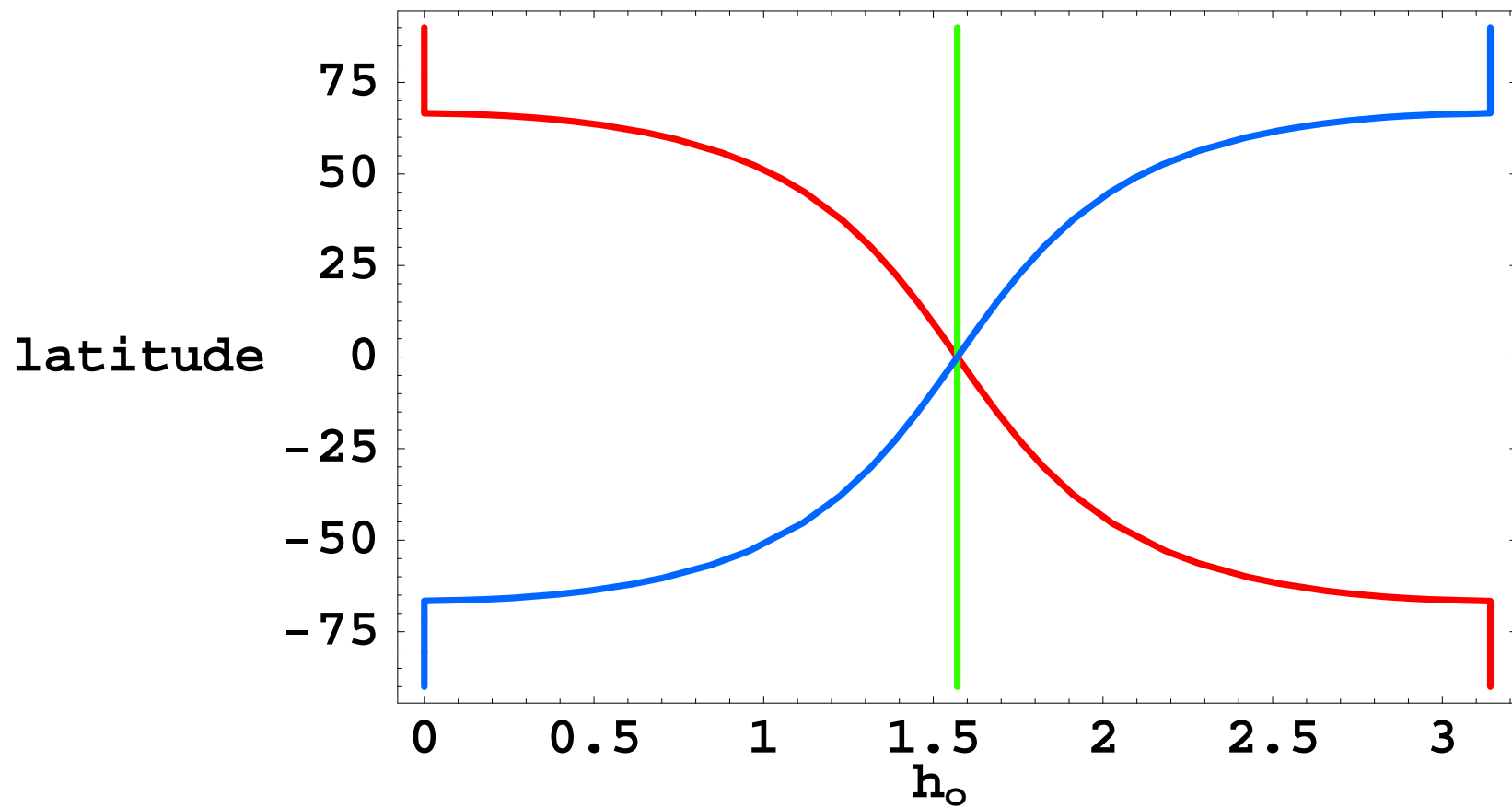
where

$$\sin \phi \sin \delta + \cos \phi \cos \delta \cos h_o = 0$$

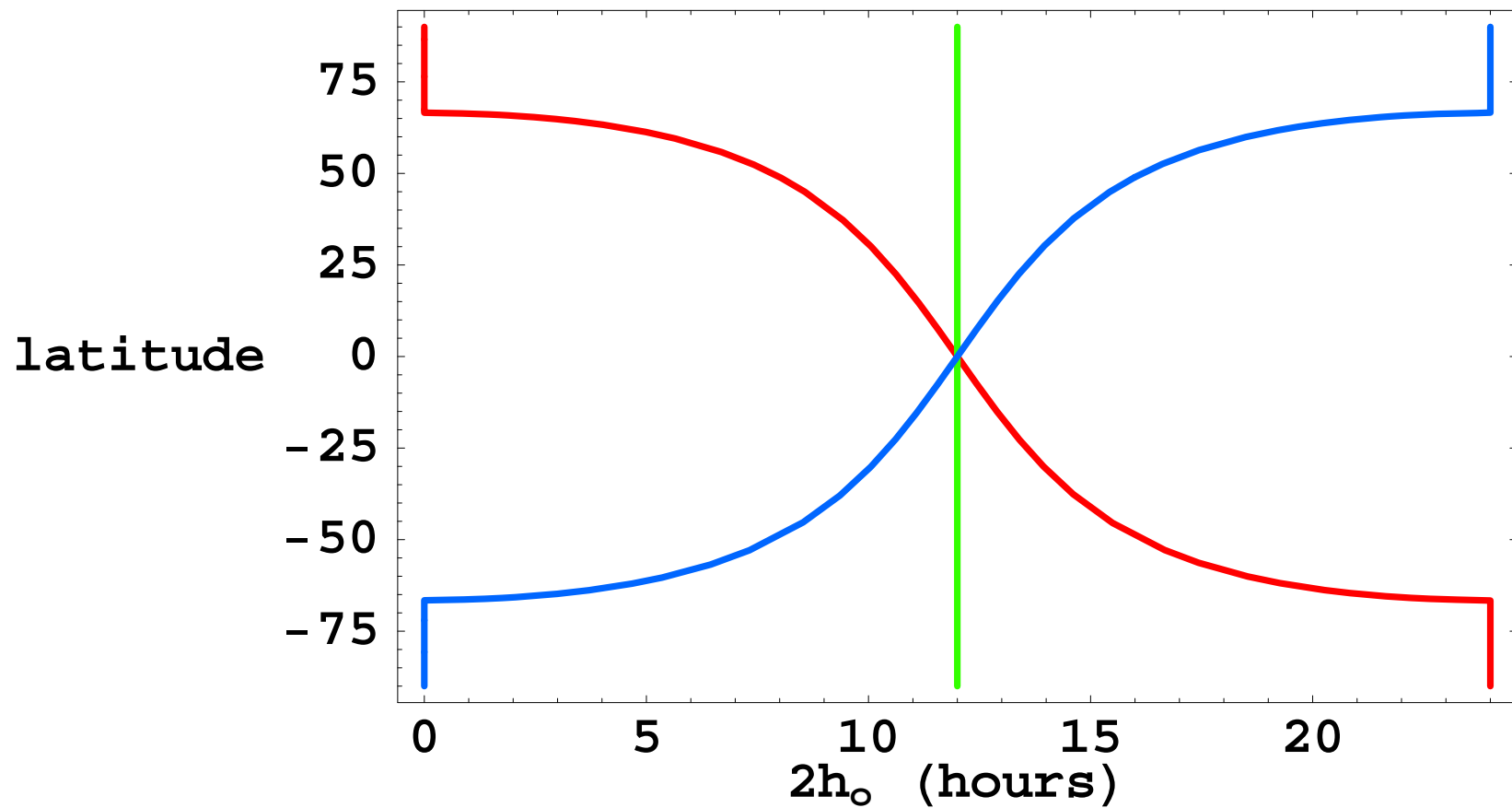
(But sometimes  $\theta_s$  is never 0 during a day, and  $h_o = \pi$  or  $h_o = 0$ .)

Keep it simple for a first calculation of  $\overline{Q}^{day}$ :

1. Assume  $\frac{R_o^2}{R_E^2} = 1$  exactly.
2. Know the current obliquity of Earth is  $23.45^\circ$ .  $\delta$  oscillates between  $\pm 23.45^\circ$ .
3. Calculate  $\overline{Q}^{day}$  at the solstices and equinoxes.

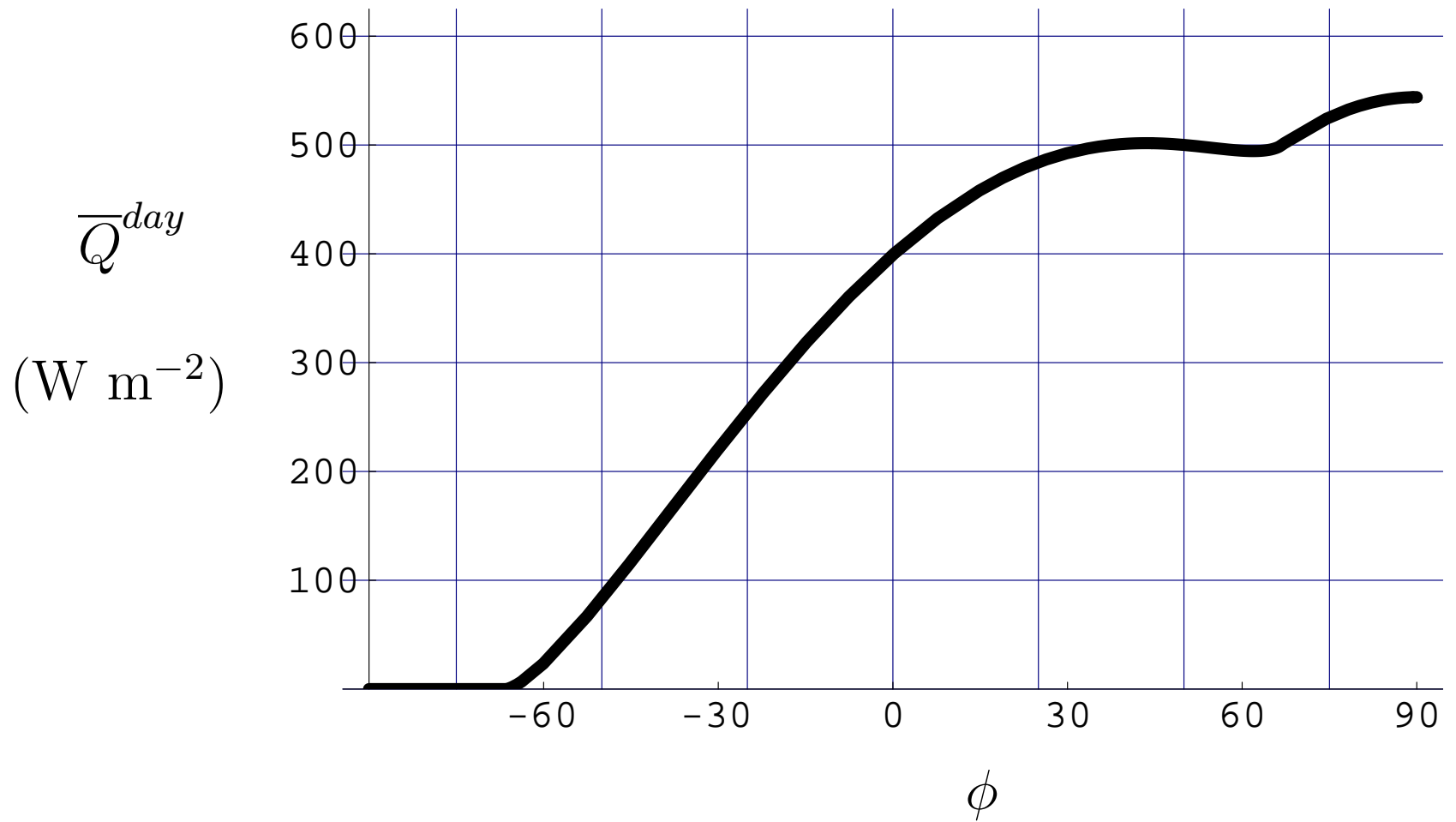


summer solstice, equinox, winter solstice



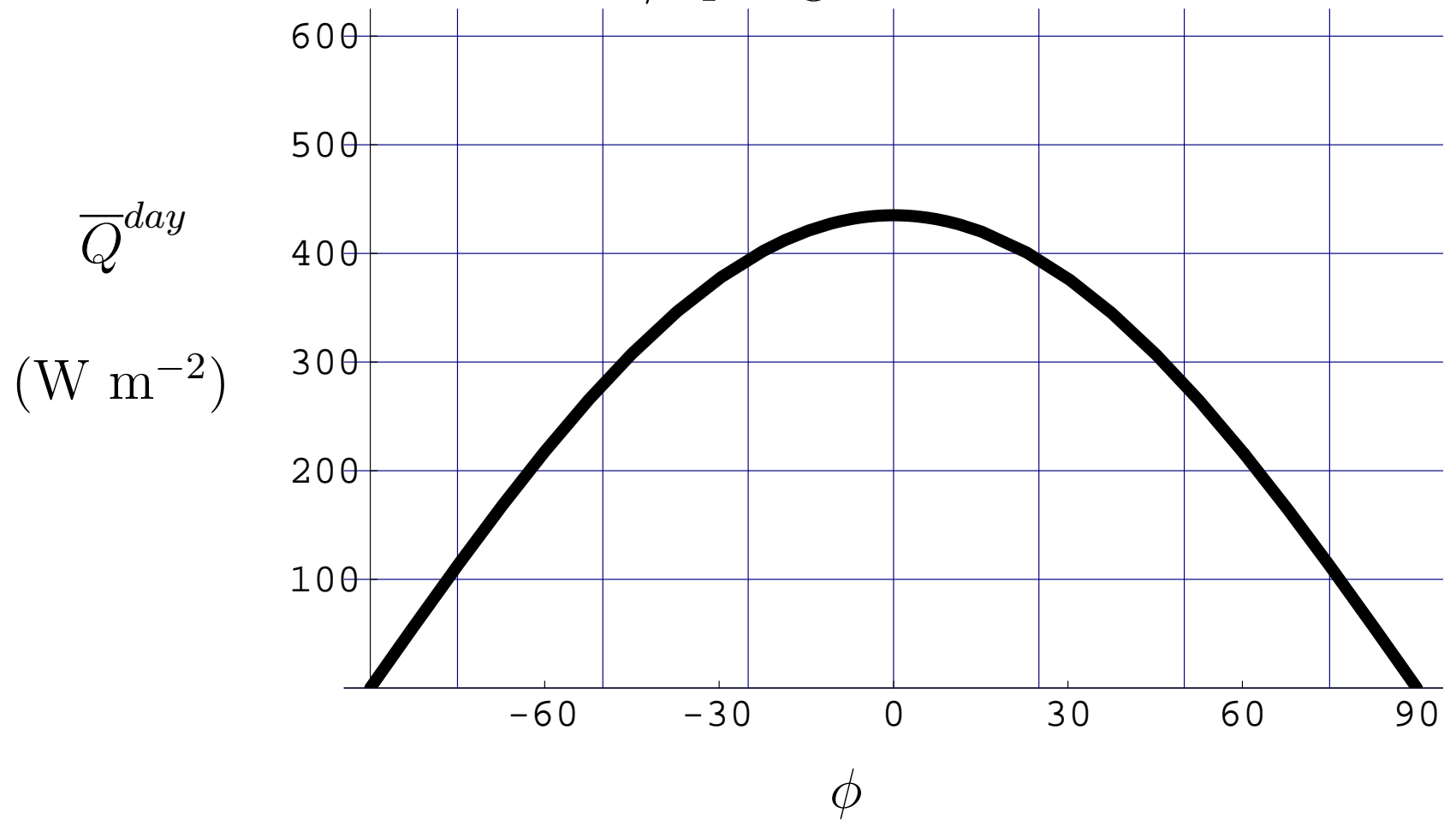
summer solstice, equinox, winter solstice

N.H. Summer  $\delta = 23.45^\circ$



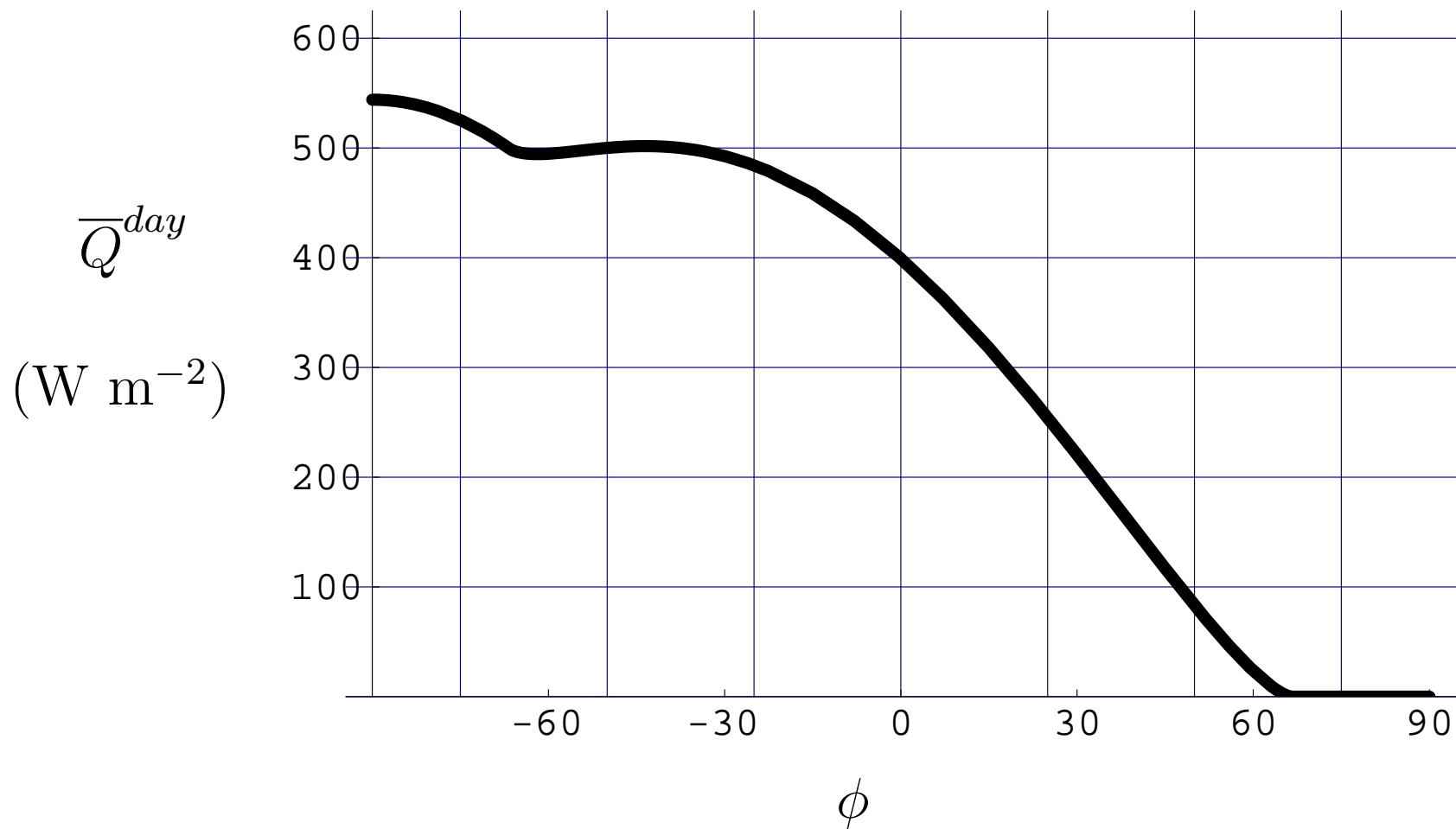
N.H. Fall/Spring

$\delta = 0^\circ$

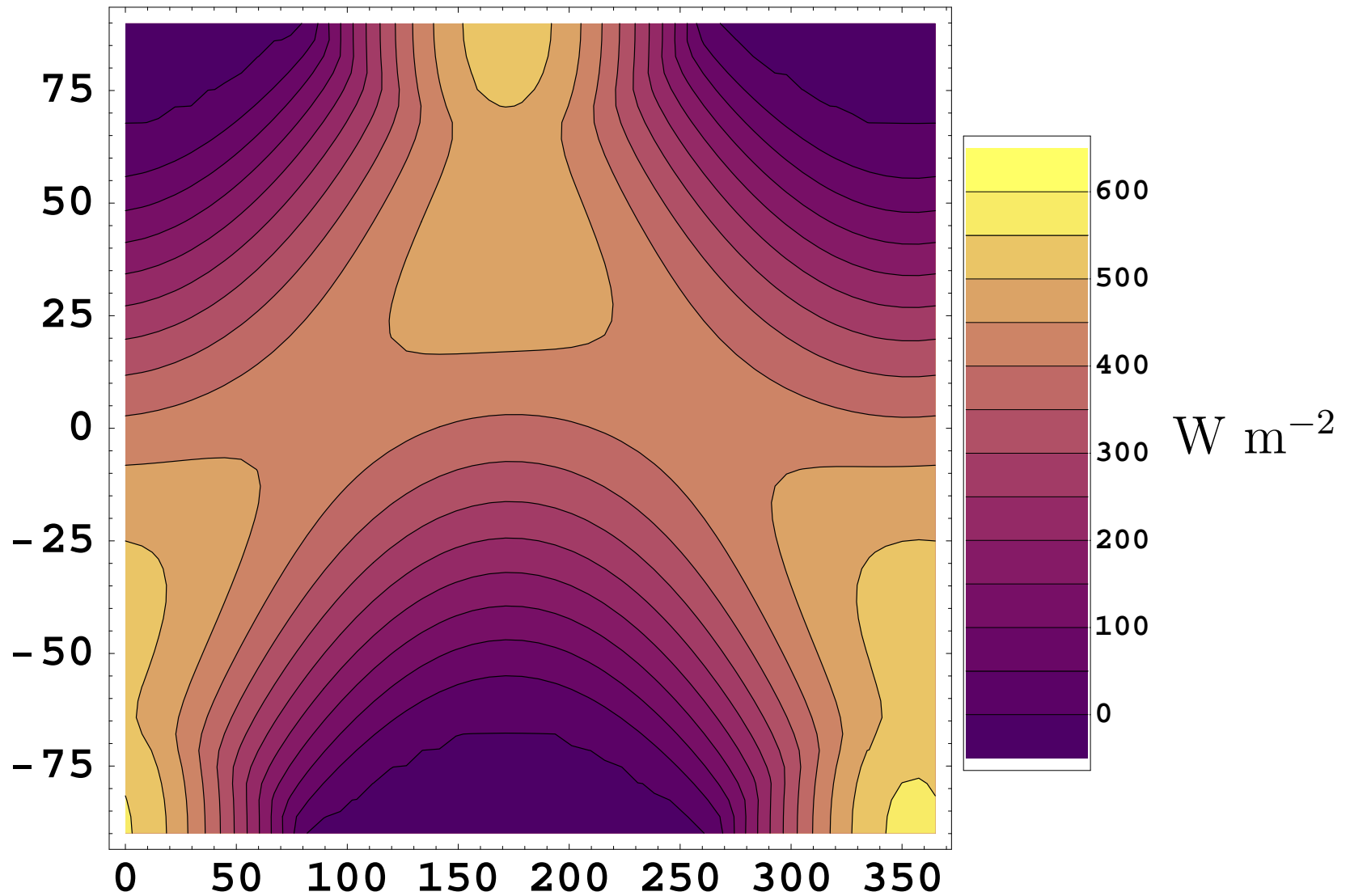


N.H. Winter

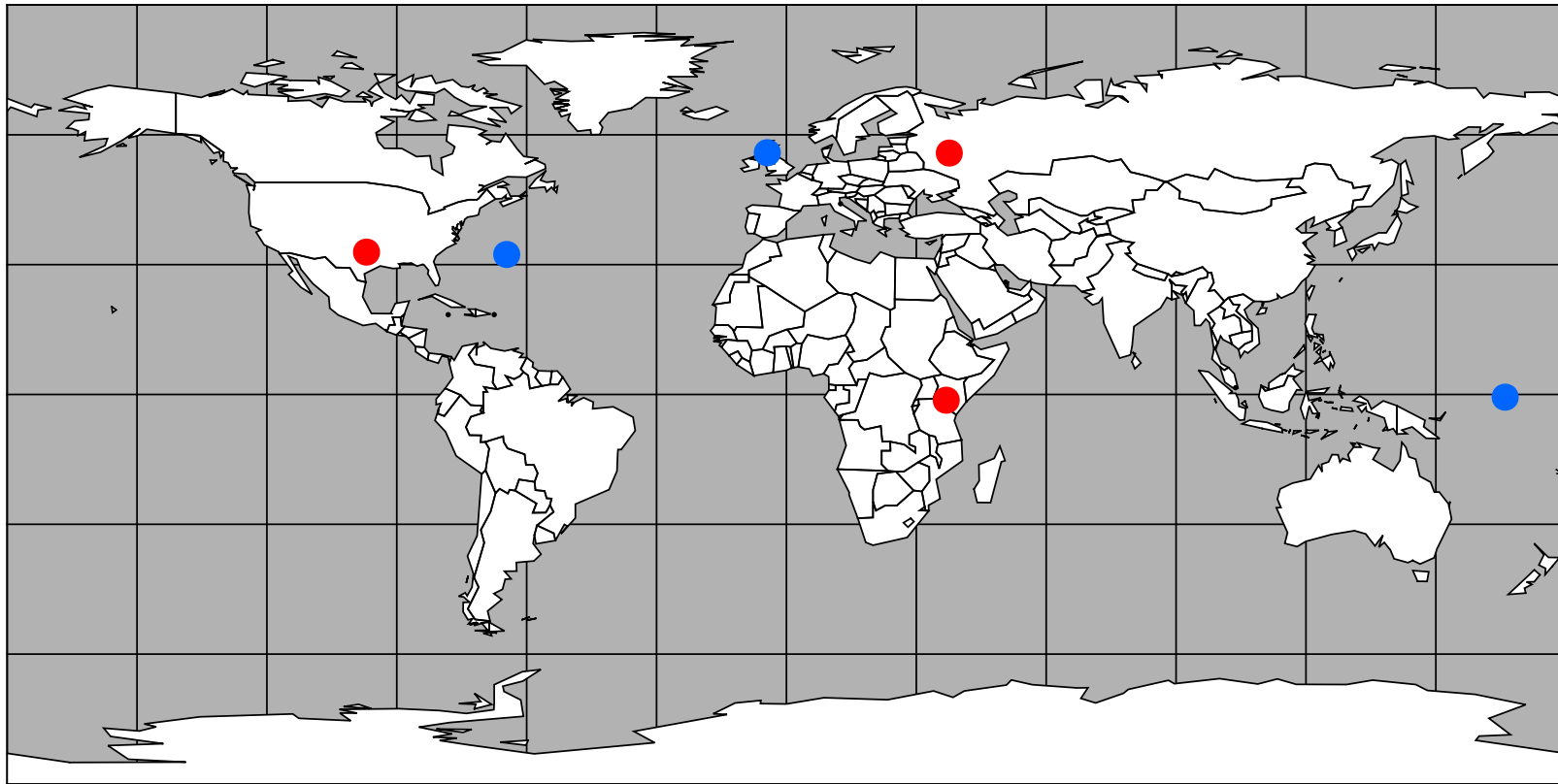
$\delta = -23.45^\circ$



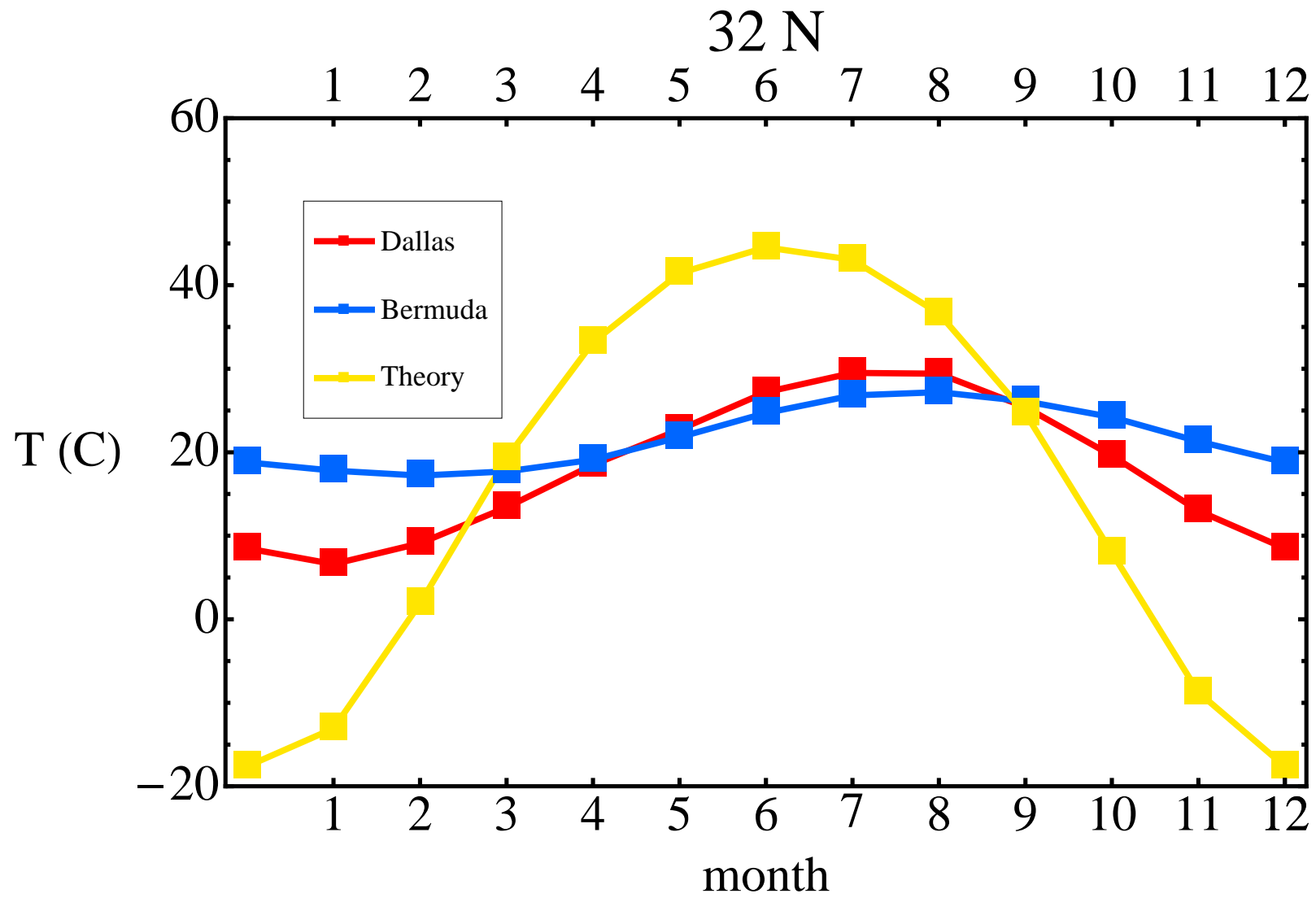
$\overline{Q}^{day}$  for 2000 A.D.



Using knowledge of how  $\frac{R_o^2}{R_E^2}$  and  $\delta$  vary with day of year.



Annual temperature cycles will be plotted at the selected sites.  
What do you anticipate?

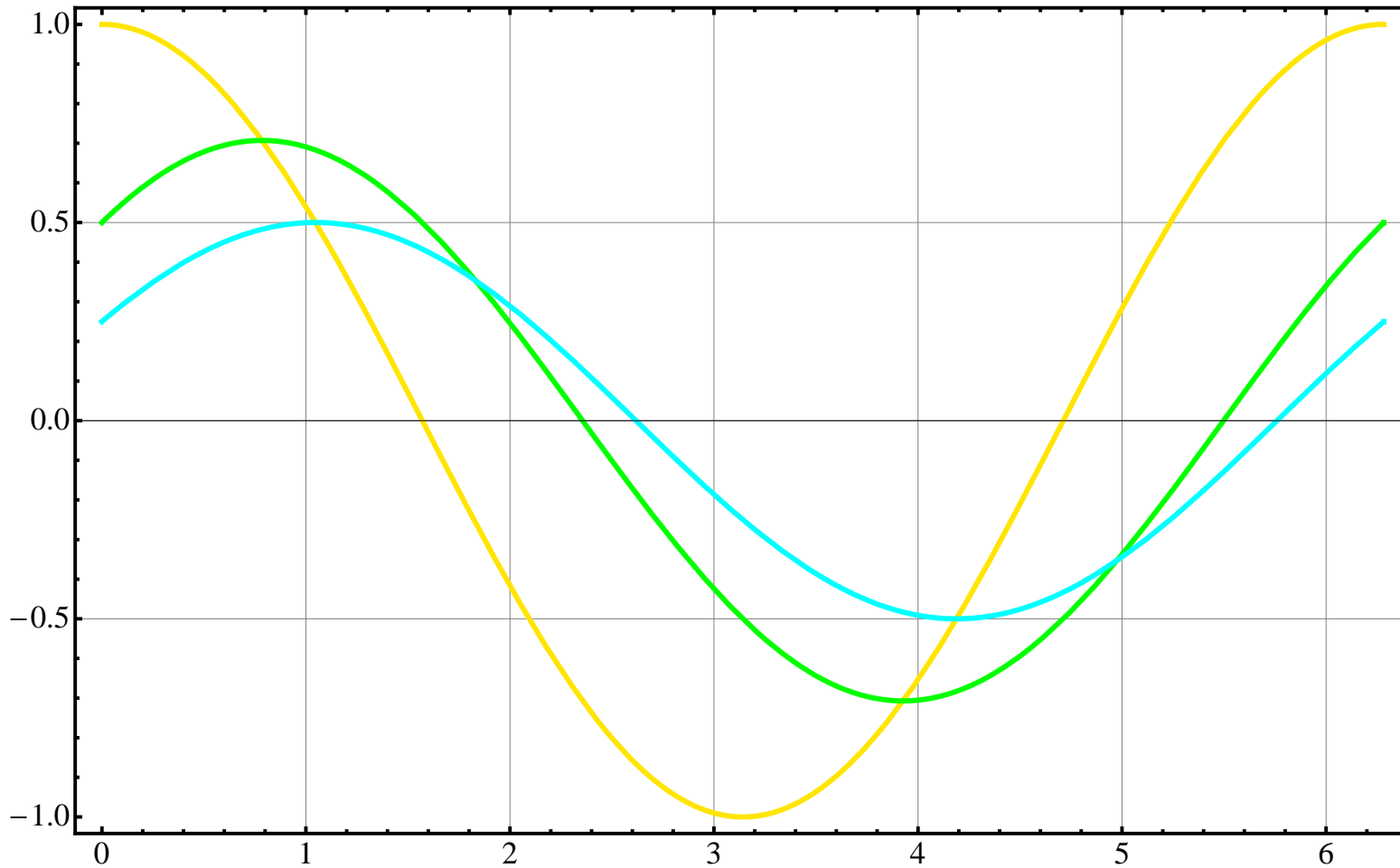


The theoretical curve is derived from our slab model:

$$\sigma T^4 = \overline{Q}^{day} \frac{1 - \alpha_p}{1 - \frac{\epsilon}{2}}$$

with  $\alpha_p = 0.3$  and  $\epsilon = 0.8$ .

The model omits thermal inertia (among other things).  
Recall the simple result that for phase lag  $\delta$  the  
amplitude is reduced by  $\cos \delta$ :



56 N

1 2 3 4 5 6 7 8 9 10 11 12

